## Contributed Talk

## Splinter CCAT

## HIGH-REDSHIFT GALAXY SURVEYS WITH CCAT-P

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Characterising the far-infrared emission of high-redshift galaxies is essential to estimate their dust-enshrouded star-formation activity otherwise only measured using not-well-understood rest-frame UV corrections. To date, far-infrared measurements are mostly based on *Herschel* or ALMA observations. However, while *Herschel* could map large areas of the sky, this 3.5 meter telescope was significantly limited by confusion (inability to separate emission from multiple galaxies within a single resolution element), providing far-infrared measurements only for the most luminous and thus not representative high-redshift galaxies.

On the other hand, ALMA can perform ultra-deep far-infrared observations but its relatively narrow field-of-view does not provide us with the statistic needed to study the evolution over cosmic time of galaxy populations.

I will demonstrate how CCAT-p, with its wide-field-of-view and 6 meter aperture, will fill this gap, providing deeper (at z>1) and wider extragalactic surveys than those from the Herschel Space Observatory. The planned CCAT-p Galaxy Evolution (GEvo) survey will improve our understanding of the faint galaxy population that is responsible for the bulk of the cosmic far-infrared background. This wide-area survey will also allow for the detection of exotic, highly luminous galaxies.